# Multi-photon decays of the <sup>7</sup> Higgs boson at the LHC <sup>7</sup> Samuel D. Lane KAIST

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arXiv: <u>2305.00013</u>



- Introduction and Motivation
- Model Overview
  - Particle content, key ideas, constraints, assumptions
- Multi-photon objects & Signal Categories
  - Photon jets, isolated photons, simulation
- Results
- Conclusion

#### **The Standard Model**

The SM works incredibly well

• The Higgs mechanism provides masses for all particles in the SM

• Nearly every measurement agrees with SM prediction



#### **The Standard Model**



#### **The Standard Model**



# What is the SM missing?



- Dark Matter
- Dark Energy
- Matter-antimatter asymmetry
  - Scale differences in the model (hierarchy, naturalness, etc.)

#### **Portals & Dark Sector**



#### **Dark Axion Portal**

- Connect ALP and dark photon
- Dark higgs, Dark photon, ALP, VLFs
- ALPs/DP could give the right relic density for dark matter
- Does not solve strong CP problem



K. Kaneta, H.-S. Lee, and S. Yun, arXiv: <u>1611.01466</u>



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#### **Particle Content**

	SU(3)	$SU(2)_L$	Y	$Y_D$	$Y_{\rm PQ}$
$\psi_L$	1	1	$Q_L$	$D_L$	$PQ_L$
$\psi_R$	1	1	$Q_R$	$D_R$	$PQ_R$
$H_{\rm PQ}$	1	1	0	0	$PQ_H$
$H_D$	1	1	0	$D_H$	0

- Global PQ symmetry, gauged dark symmetry
- SM particles have no PQ or dark charge
- BSM Charges chosen to cancel anomalies

#### **Kinetic Mixing**



- Experimental limits are important
- The VLFs induce kinetic mixing
- For O(1) dark gauge couplings expect O(0.1) kinetic mixing
- We will assume zero kinetic mixing

# Tracking the Goldstone modes

# Normal SM Higgs mechanism

- SM Higgs goldstones become the longitudinal modes of W and Z
- Dark Higgs mechanism
  - H<sub>D</sub> goldstone becomes longitudinal modes of dark photon
- PQ mechanism at some scale f<sub>a</sub>
  - H<sub>PQ</sub> goldstone becomes the ALP

### **SM Higgs Mechanism**



$$V(\varphi) = -\mu^2 (\varphi^{\dagger} \varphi) + \lambda (\varphi^{\dagger} \varphi)^2.$$
$$\frac{\partial \mathcal{V}}{\partial |\varphi|^2} = \mu^2 - 2\lambda \varphi_0^{\dagger} \varphi_0 = 0$$
$$\varphi_0^{\dagger} \varphi_0 = v^2 = \frac{\mu^2}{2\lambda}$$
$$\varphi = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2}G^+ \\ h + v + iG_0 \end{pmatrix}$$

#### **Scalar Sector Summary**

- Potential depends on all scalars
- Should in principle write down the most general renormalizable potential
- Each scalar acquires a vev
- Scalar kinetic terms give masses to vetor bosons
- There could be some scalar mixing between the CP even neutral states

 $\mathcal{V}(\Phi, H_D, H_{\mathrm{PQ}})$ 

$$|D_{\mu}\Phi|^2 + |D_{\mu}H_D|^2$$

$$\binom{h_1}{h_2} = \hat{R}(\theta_S) \binom{h}{h_d}$$

# **Higgs Couplings**

• Higgs dark photon coupling



• In principle there is a Higgs ALP coupling

$$|D_{\mu}H_{D}|^{2} \longrightarrow \frac{\lambda_{h\gamma_{D}\gamma_{D}}}{2}h\gamma_{D}^{\mu}\gamma_{D\mu}$$

$$H_{D} = \frac{1}{\sqrt{2}}(h_{D} + v_{D})$$

$$\lambda_{h\gamma_{D}\gamma_{D}} \sim \sin\theta \frac{m_{\gamma_{D}}^{2}}{v_{D}}.$$

$$h \cdots \int VLF$$

$$u = a$$

# **ALP Couplings**

$$\frac{G_{a\gamma\gamma}}{4}aF^{\mu\nu}\tilde{F}_{\mu\nu} + \frac{G_{a\gamma\gamma_D}}{2}aF_D^{\mu\nu}\tilde{F}_{\mu\nu} + \frac{G_{a\gamma_D\gamma_D}}{4}aF_D^{\mu\nu}\tilde{F}_{D\mu\nu}$$

- ALP-photon-photon
- ALP-photon-dark photon
- ALP-dark photon-dark photon



$$G_{a\gamma\gamma} = \frac{e^2}{8\pi^2} \frac{PQ_{\Phi}}{f_a} \Big[ 2N_C Q_{\psi}^2 - \frac{2}{3} \frac{4+z}{1+z} \Big],$$
  

$$G_{a\gamma\gamma'} \simeq \frac{ee'}{8\pi^2} \frac{PQ_{\Phi}}{f_a} \Big[ 2N_C D_{\psi} Q_{\psi} \Big] + \varepsilon G_{a\gamma\gamma},$$
  

$$G_{a\gamma'\gamma'} \simeq \frac{e'^2}{8\pi^2} \frac{PQ_{\Phi}}{f_a} \Big[ 2N_C D_{\psi}^2 \Big] + 2\varepsilon G_{a\gamma\gamma'}.$$



# Why photons?

- Axion to diphoton is well known
- Higgs to diphoton is also well known
- Photons are "clean" at colliders
- Go look for additional signals at LHC that contain photons





# **Some Signals With Photons**

• Some candidate signals in the dark axion portal



#### **Higgs to six photons**

 $\sigma(pp \to h \to \gamma_D \gamma_D \to 6\gamma) \approx \sigma(pp \to h) BR (h \to \gamma_D \gamma_D) BR^2 (\gamma_D \to a\gamma) BR^2 (a \to \gamma\gamma)$ 

- Assume Shell decays
- Use narrow width approximation
- Take the spin averaged matrix element





#### **Full Branching Ratio**

$$BR(h \to \gamma_D(k_1)\gamma_D(k_2) \to \gamma(p_1)a(q_1)\gamma(p_2)\gamma(q_2) \to \gamma(p_1)\gamma(p_3)\gamma(p_4)\gamma(p_2)\gamma(p_5)\gamma(p_6))$$
  
= BR(h \to \gamma\_D\gamma\_D) BR(\gamma\_D \to a\gamma)^2 BR(a \to \gamma\gamma)^2  $\frac{9\int d\Omega_h d\Omega_{\gamma_D(k_1)} d\Omega_{\gamma_D(k_2)}f(p_1, k_1, p_2, k_2)}{(2\pi)^3 \left(12 - 4\frac{m_h^2}{m_{\gamma_D}^2} + \frac{m_h^4}{m_{\gamma_D}^4}\right) (m_{\gamma_D}^2 - m_a^2)^4}$ 

 $f(p_1, k_1, p_2, k_2) = (p_2 \cdot k_2)(k_1 \cdot k_2)(p_1 \cdot p_2)(p_1 \cdot k_1) + (p_2 \cdot k_1)(p_2 \cdot k_2)(p_1 \cdot k_2)(p_1 \cdot k_1)$  $- (p_2 \cdot k_1)(p_1 \cdot p_2)(p_1 \cdot k_1)k_2^2 - (p_2 \cdot k_2)(p_1 \cdot k_2)(p_1 \cdot p_2)k_1^2$  $+ \frac{1}{2}(p_1 \cdot p_2)^2k_1^2k_2^2.$ 

### **Higgs Production**



(21)

### **Higgs Constraints**

95% CL



#### **Axion Constraints**

ALP-photon-dark photon

ALP-photon-photon



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#### **Average Decay Lengths**

Average Minimum Dark Photon Decay Length Only  $\gamma_D \rightarrow a\gamma$  with  $G_{a\gamma\gamma_D} = 0.002 \text{ GeV}^{-1}$ 10<sup>1</sup>  $m_{\gamma_D}$  (GeV)  $1 \,\mu m$ 10<sup>0</sup> 100 µm 1 mm 1 cm .0 cm  $10^{-1}$  $10^{-1}$  $10^{-2}$  $10^{0}$  $10^{1}$  $m_a$  (GeV)

$$\Gamma(\gamma_D \to \gamma a) = \frac{|G_{a\gamma\gamma_D}|^2 m_{\gamma_D}^3}{96\pi} \left(1 - \frac{m_a^2}{m_{\gamma_D}^2}\right)^3$$

$$G_{a\gamma\gamma_D} \approx 2 \times 10^{-3} \text{ GeV}^{-1}$$

Mostly prompt decays with some displaced vertices



#### **Average Decay Lengths**

$$\Gamma(a o \gamma \gamma) = rac{|G_{a\gamma\gamma}|^2 m_a^3}{64\pi}$$
  
Maximal  $G_{a\gamma\gamma}$ 

10<sup>1</sup>

10<sup>0</sup>

Mostly prompt decays, with some displaced vertices, and some missing energy 10-1

Only  $a \rightarrow \gamma \gamma$  with maximum  $G_{a\gamma\gamma}$  coupling Р hun 100 *µ*m 1 mm щ cm  $10^{-1}$  $10^{-2}$  $10^{1}$  $10^{0}$  $m_a$  (GeV)

Average Minimum ALP Decay Length

#### **Model Assumptions**

- Want on-shell decays
- Zero kinetic mixing
- Dark photon only decays to ALP and photon
- ALP only decays to photon pairs

$$m_h > m_{\gamma_D} > m_a$$
  $\varepsilon \to 0$   
 $BR(\gamma_D \to a\gamma) = 1$   $BR(a \to \gamma\gamma) = 1$ 





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#### **Photon Jets**



#### **Multi-Photon Objects**

$$\Delta R = \sqrt{\Delta \eta^2 + \Delta \phi^2}$$

$$\eta = \frac{1}{2} \ln \frac{E + p_Z}{E - p_Z}$$

Well collimated photons end up in same detector location

Appear as a single photon  $\Delta R < 0.04$ 

**Photon Jets** 

Sets of photons or photon-jets that have intermediate separation

 $0.04 < \Delta R < 0.4$  **ξ Jets** 

**B. Sheff et al.** 2008.10568

Use Isolated Photons to reduce QCD backgrounds

 $\Delta R > 0.4$  Isolated Photons

#### **Generate Truth Level Events**

$$h \to \gamma_d \gamma_d \to a \gamma a \gamma \to 6 \gamma$$

1. Truth level events  $\{\gamma\}$ 

- Parton Luminoisty
- Spin-correlations  $\mathcal{L}(\tau) =$
- Uniformly sample phase space
- Uniformly sample higgs rapidity

$$\tau = m_h^2/S, \ \sqrt{S} = 13 \text{ TeV}$$
$$\int_{\ln\sqrt{\tau}}^{-\ln\sqrt{\tau}} dy_h g(\sqrt{\tau}e^{y_h}, \mu_F) g(\sqrt{\tau}e^{-y_h}, \mu_F),$$

$$\frac{9\int d\Omega_h d\Omega_{\gamma_D(k_1)} d\Omega_{\gamma_D(k_2)} f(p_1, k_1, p_2, k_2)}{(2\pi)^3 \left(12 - 4\frac{m_h^2}{m_{\gamma_D}^2} + \frac{m_h^4}{m_{\gamma_D}^4}\right) (m_{\gamma_D}^2 - m_a^2)^4}$$



### **Merge Photon Jets**

$$h \to \gamma_{d}\gamma_{d} \to a\gamma a\gamma \to 6\gamma \to n\gamma$$
1. Truth level events { $\gamma$ }  
Merge  $\gamma$ -Jets  $\Delta R < 0.04$   
2. Observable Photons { $\gamma_{obs}$ }  

$$Prob(6\gamma \to n\gamma_{iso} + m\xi) = \frac{BR(h \to \gamma_{D}\gamma_{D} \to a\gamma a\gamma \to 6\gamma \to n\gamma_{iso} + m\xi)}{BR(h \to \gamma_{D}\gamma_{D})BR^{2}(\gamma_{D} \to a\gamma)BR^{2}(a \to \gamma\gamma)} Max Prob(6\gamma \to n\gamma_{obs})$$

$$Max Prob(6\gamma \to n\gamma_{obs})$$

#### **Signals after Merging**



# **Estimating Trigger/Detector Efficiency**

- Use simple cut and count
- Impose some transverse momentum cuts on k photons
- Impose some rapidity cuts on all photons
- Try to match existing searches when possible

$$\operatorname{Eff}_{k}(h \to n\gamma_{iso} + m\xi)$$





### $|\eta| < 1.44$ or $1.57 < |\eta| < 2.5$ .

Channel	CMS $p_T$ Requirements
$1\gamma$	$p_{1,T} > 145 \text{ GeV} [98]$
$2\gamma$	$p_{1,T} > 30 \text{ GeV} \text{ and } p_{2,T} > 18 \text{ GeV} [30]$
$3\gamma$	$p_{1,T} > 15 \text{ GeV}, p_{2,T} > 15 \text{ GeV}, \text{ and } p_{3,T} > 15 \text{ GeV}$ [95]
$4\gamma$	$p_{1,T} > 30 \text{ GeV}, p_{2,T} > 18 \text{ GeV}, p_{3,T} > 15 \text{ GeV}, \text{ and } p_{4,T} > 15 \text{ GeV}$ [27]
$5\gamma$	$p_{i,T} > 15 \text{ GeV} (i = 1, 2, 3, 4, 5)$
$6\gamma$	$p_{i,T} > 15 \text{ GeV} (i = 1, 2, 3, 4, 5, 6)$

#### **ATLAS Cuts**

#### $|\eta| < 1.37$ or $1.52 < |\eta| < 2.5$

Channel	ATLAS $p_T$ Requirements
$1\gamma$	$p_{1,T} > 150 \text{ GeV} [94]$
$2\gamma$	$p_{1,T} > 35 \text{ GeV} \text{ and } p_{2,T} > 25 \text{ GeV} [24]$
$3\gamma$	$p_{1,T} > 15 \text{ GeV}, p_{2,T} > 15 \text{GeV}, \text{ and } p_{3,T} > 15 \text{ GeV}$ [95]
$4\gamma$	$p_{1,T} > 30 \text{ GeV}, p_{2,T} > 18 \text{ GeV}, p_{3,T} > 15 \text{ GeV}, \text{ and } p_{4,T} > 15 \text{ GeV}$ [95]
$5\gamma$	$p_{i,T} > 15 \text{ GeV} (i = 1, 2, 3, 4, 5)$
$6\gamma$	$p_{i,T} > 15 \text{ GeV} (i = 1, 2, 3, 4, 5, 6)$



#### **Estimated Trigger Efficiencies**



#### **Estimated Trigger Efficiencies**





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#### **Two Photon Signals**



#### **Four Photon Signals**



### **Six Photon Signals**











#### **Combining Constraints**





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#### **Summary & Conclusion**

- The DAP introduces a six photon Higgs resonance.
- We can place good constraints using the two and four photon categories.
- Could constrain other regions by doing appropriate searches
- The pure six photon signal has a good chance to be seen



# Thanks for your attention!

# Questions?

