# Scattering Amplitudes and Goldstone Boson Equivalence Theorems In Extra Dimensions

# Dipan Sengupta

# With R. Sekhar Chivukula (UCSD), Dennis Foren (UCSD), Kirtimaan A. Mohan (MSU) , Elizabeth H. Simmons(UCSD), X. Wang (UCSD) and J. A. Gill (Adelaide)

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Brane Localized Scalar Dark Matter annihilating via a KK/massive spin-2 portal



# How calculations go wrong in DM

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Gravity-mediated Scalar Dark Matter in Warped Extra-Dimensions

PRL 116, 101302 (2016)

PHYSICAL REVIEW LETTERS

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Mathias Garny,<sup>1,\*</sup> McCullen Sandora,<sup>2,†</sup> and Martin S. Sloth<sup>2,‡</sup> <sup>1</sup>CERN Theory Division, CH-1211 Geneva 23, Switzerland  $^{2}CP^{3}$ -Origins, Center for Cosmology and Particle Physics Phenomenology, University of Southern Denmark, Campusvej 55, 5230 Odense M, Denmark (Received 20 November 2015; published 10 March 2016)

Miguel G. Folgado,<sup>a</sup> Andrea Donini,<sup>a</sup> Nuria Rius<sup>a</sup>

### Blatantly wrong-60 Citations

Spin-2 portal dark matter

Nicolás Bernal,<sup>1,\*</sup> Maíra Dutra,<sup>2,†</sup> Yann Mambrini,<sup>2,‡</sup> Keith Olive,<sup>3,4,§</sup> Marco Peloso,<sup>3,||</sup> and Mathias Pierre<sup>2,¶</sup> <sup>1</sup>Centro de Investigaciones, Universidad Antonio Nariño Carrera 3 Este # 47A-15, Bogotá, Colombia <sup>2</sup>Laboratoire de Physique Théorique (UMR8627), CNRS, University of Paris-Sud, Université Paris-Saclay, 91405 Orsay, France

Partially wrong-100 Citations

# Partially wrong ~ 150 Citations

Dark Matter Direct Detection from new interactions in models with spin-two mediators

A. Carrillo-Monteverde<sup>1, $\sharp$ </sup>, Yoo-Jin Kang<sup>2,3,\*</sup>, Hyun Min Lee<sup>2, $\dagger$ </sup>, Myeonghun Park<sup>4, $\bigstar$ </sup>, and Veronica Sanz<sup>1, $\bigstar$ </sup>

Department of Physics and Astronomy, University of Sussex, Brighton BN1 9QH, UK <sup>2</sup>Department of Physics, Chung-Ang University, Seoul 06974, Korea. <sup>3</sup>Center for Theoretical Physics of the Universe, Institute for Basic Science, Daejeon, 34051, Korea. <sup>4</sup>Institute of Convergence Fundamental Studies and School of Liberal Arts, Seoul National University of Science and Technology, Seoul 01811, Korea



week ending 11 MARCH 2016

### PHYSICAL REVIEW LETTERS **128**, 081806 (2022)

### Massive Gravitons as Feebly Interacting Dark Matter Candidates

Haiying Cai<sup>(0)</sup>,<sup>1,\*</sup> Giacomo Cacciapaglia<sup>(0)</sup>,<sup>2,3,†</sup> and Seung J. Lee<sup>(1,‡</sup> <sup>1</sup>Department of Physics, Korea University, Seoul 136-713, Korea <sup>2</sup>University of Lyon, Université Claude Bernard Lyon 1, F-69001 Lyon, France <sup>3</sup>Institut de Physique des 2 Infinis de Lyon (IP2I), UMR5822, CNRS/IN2P3, F-69622 Villeurbanne Cedex, Franc

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### Partially wrong ~ 150 Citations

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### Wrong Scaling

 $||\mathcal{M}|^2 \sim E^{12}/(m^8 m_p^4)|$ 

е			

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Issue: Breakdown of Diffeomorphism in naive massive gravity is restored in a full Kaluza-Klein theory

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# Massive Gauge Theories

### Massive Gauge Theories



$$-\frac{1}{2}m^2A_\mu A^\mu$$

### Mass term explicitly breaks the gauge redundancy

$$\frac{1}{2}m^2 A_{\mu}A^{\mu} - mA_{\mu}\partial^{\mu}\phi - \frac{1}{2}\partial_{\mu}\phi\partial^{\mu}\phi + A_{\mu}J^{\mu} - \frac{1}{m}\phi\partial_{\mu}J^{\mu} \qquad \delta\phi =$$





# Massive Gauge Theories

Let's think of a Massive Photon 
$$\mathcal{L}_{Proce} = -\frac{1}{4}F_{\mu\nu}^{2} - \frac{1}{2}m^{2}A_{\mu}A^{\mu}$$
 Mass term explicitly breaks the gauge reduped of the Higgs 
$$\mathcal{L}_{Proce} = -\frac{1}{4}F_{\mu\nu}^{2} - \frac{1}{2}m^{2}A_{\mu}A^{\mu} - mA_{\mu}\partial^{\mu}\phi - \frac{1}{2}\partial_{\mu}\partial\partial^{\mu}\phi + A_{\mu}J^{\mu} - \frac{1}{m}\phi\partial_{\mu}J^{\mu}$$

$$\frac{\delta A_{\mu} - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2}m^{2}A_{\mu}A^{\mu} - mA_{\mu}\partial^{\mu}\phi - \frac{1}{2}\partial_{\mu}\partial\partial^{\mu}\phi + A_{\mu}J^{\mu} - \frac{1}{m}\phi\partial_{\mu}J^{\mu}$$

$$\frac{\delta A_{\mu} - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2}m^{2}A_{\mu}A^{\mu} - mA_{\mu}\partial^{\mu}\phi - \frac{1}{2}\partial_{\mu}\partial\partial^{\mu}\phi + A_{\mu}J^{\mu} - \frac{1}{m}\phi\partial_{\mu}J^{\mu}$$

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$$\frac{\delta A_{\mu} - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2}(\mathcal{D}_{\mu}\phi)(\mathcal{D}^{\mu}\phi)^{*} - \lambda(\phi\phi^{*} - \Phi_{0}^{2})^{2}$$

$$\frac{\phi = \varphi e^{i\chi}}{\phi}$$

$$\text{VEV of the Higgs}$$

$$\frac{\mathcal{L}_{A\Pi} = -\frac{1}{4}F_{\mu\nu}^{2} - \frac{1}{2}\varphi^{2}(qA_{\mu} - \partial_{\mu}\chi)^{2} - \frac{1}{2}(\partial_{\mu}\varphi)^{2} - \lambda(\varphi^{2} - \Phi_{0}^{2})^{2}$$

$$\frac{A_{\mu} \rightarrow A_{\mu} + \partial_{\mu}\xi(x)}{\chi \rightarrow \chi + q\xi(x)}$$
Restoration of U(1)





### Goldstone Boson Equivalence Theorems for the Standard Model W's





Courtesy Michael Peskin and Kurt Gottfried

 $\begin{aligned} & \text{Polarization vectors of massive W's} \quad k^{\mu} = (E_k, 0, 0, k) \\ & \epsilon_L^{\mu}(k) = \left(\frac{k}{m_W}, 0, 0, \frac{E_k}{m_W}\right) = \frac{k^{\mu}}{m_W} + \frac{1}{m_W} \left(-E_k + k, 0, 0, E_k - k\right) \\ & E_k \geq m_W \quad = \frac{k^{\mu}}{m_W} + \mathcal{O}\left(\frac{m_W}{E_k}\right) \end{aligned}$ 

Polarization vectors increasingly parellel to momentum vector







# Goldstone Boson Equivalence Theorems for the Standard Model W's

$$\mathcal{M}_{4W} + \mathcal{M}_{\gamma}^{s} + \mathcal{M}_{\gamma}^{t} + \mathcal{M}_{Z}^{s} \mathcal{M}_{Z}^{t} = \frac{g^{2}(s+t)}{4m_{W}^{2}} - \frac{g^{2}(s^{2}+st-t)}{2(st-t)^{2}} - \frac{g^{2}(st-t)}{2(st-t)^{2}} - \frac{g^{2}(st-t)}{2(st-$$

Adding the two reduces the growth to a constant and restores unitarity



 $4t^2\cos^2\theta_W + t^2)$ Improvement from E<sup>4</sup> to E<sup>2</sup>  $\cos^2 \theta_W$ )

$$\left|\mathcal{M}_{(W_L^- W_L^+ \to W_L^- W_L^+)} = \frac{g^2}{2} \left(-\frac{m_H^2}{m_W^2} - \frac{s^2 + st - st}{st \cos^2 \theta}\right)\right|$$

Using the Goldstone Equivalence Theorem : In 't-Hooft Feynman Gauge replace longitudinal W's with Goldstones  $\,w^{\pm}$ 

No Large Cancellations that happened at low energies





### The Classical Action for Gravity



**Conserved Source** 

 $\left|R_{\mu\nu} - \frac{\mathbf{I}}{2}g_{\mu\nu}R = -8\pi G T_{\mu\nu}\right|$ 

 $\sqrt{g}T^{\mu\nu} \equiv -2\frac{\partial}{\partial g_{\mu\nu}}\left(\sqrt{g}\mathcal{L}_m\right)$ 

### The Classical Action for Gravity



The Ricci is a two derivative object  $g_{\mu\nu} \rightarrow \eta_{\mu\nu} + \kappa h$ 

Einstein Hilbert Action is a dimension 6 operator with a cut-off MPI

**Conserved Source** 

$$R = -8\pi G T_{\mu\nu}$$

$$\sqrt{g}T^{\mu\nu} \equiv -2\frac{\partial}{\partial g_{\mu\nu}} \left(\sqrt{g}\mathcal{L}_m\right)$$

$$\dot{n}_{\mu
u}$$



 $\begin{aligned} \epsilon_{\pm 2}^{\mu\nu} &= \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu} \\ \epsilon_{\pm 1}^{\mu} &= \pm \frac{e^{\pm i\phi}}{\sqrt{2}} \left( 0, -c_{\theta} c_{\phi} \pm i s_{\phi}, -c_{\theta} s_{\phi} \mp i c_{\phi}, s_{\theta} \right) \end{aligned}$ 

### The Classical Action for Gravity



The Ricci is a two derivative object  $g_{\mu\nu} \rightarrow \eta_{\mu\nu} + \kappa h_{\mu\nu}$ 

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Consistent with semi-classical gravity

$$_{1}R^{2}+c_{2}R_{\mu\nu}R^{\mu\nu}+\ldots+\mathcal{L}_{matter}$$

 $ic_{\phi}, s_{\theta}$ 

Diffeomorphism

 $h_{\mu\nu} \to h_{\mu\nu} + \partial_{(\mu}\xi_{\nu)}$ 

Equivalent to transformation for Gauge Theories

D.O.F counting in d dimensions for the massless graviton

$$d(d+1)/2 - 2d = d(d-3)/2$$

D.O.F counting in d dimensions for the massless gauge boson



$$A_{\mu} \to A_{\mu} + \partial_{\mu}\xi$$

Diffeomorphism

$$h_{\mu\nu} \to h_{\mu\nu} + \partial_{(\mu}\xi_{\nu)}$$

Equivalent to transformation for Gauge Theories

D.O.F counting in d dimensions for the massless graviton

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D.O.F counting in d dimensions for the massless gauge boson

Let's think of a Massive Photon

$$\mathcal{L}_{\rm Proca} = -\frac{1}{4}F_{\mu\nu}^2 - \frac{1}{2}m^2 A_{\mu}A^{\mu}$$

$$\begin{array}{ll} \mbox{Propagator} & \frac{-i}{p^2 + m^2} \left( \eta_{\mu\nu} + \frac{p_{\mu}p_{\nu}}{m^2} \right) \\ \\ \mbox{Stuckelberg Trick} & A_{\mu} \rightarrow A_{\mu} + \partial_{\mu}\phi & \boxed{-\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2}m^2A_{\mu}A_{\mu}} \\ \\ \mbox{Propagators} & \frac{-i\eta_{\mu\nu}}{p^2 + m^2}, & \frac{-i}{p^2 + m^2} \end{array}$$

$$A_{\mu} \to A_{\mu} + \partial_{\mu}\xi$$

Mass term explicitly breaks the gauge redundancy

$$\delta A_{\mu} = \partial_{\mu} \Lambda$$
$$\delta A_{\mu} = \partial_{\mu} \Lambda$$
$$\delta \phi = -m\Lambda$$



Let's think of a Massive Graviton

$$S_G = \int d^4x \sqrt{g}R + m^2((h_{\mu\nu})^2 - h^2) + \kappa h_{\mu\nu}T^{\mu\nu}$$
 Fierz-Pauli Theory:



Let's think of a Massive Graviton

$$S_G = \int d^4x \sqrt{g}R + m^2((h_{\mu\nu})^2 - h^2) + \kappa h_{\mu\nu}T^{\mu\nu}$$
 Fierz-Pauli Theory

Stuckelberg

$$h_{\mu\nu} \to h_{\mu\nu} + \partial_{\mu}A_{\nu} + \partial_{\nu}A_{\mu}$$

rescale  $A_{\mu} 
ightarrow rac{1}{m} A_{\mu}, \ \phi 
ightarrow rac{1}{m^2} \phi$  Assume source is conserved , vanishing

$$S = \int d^{D}x \ \mathcal{L}_{m=0}(h') - \frac{1}{2} F_{\mu\nu} F^{\mu\nu} - 2 \frac{D-1}{D-2} \partial_{\mu} \phi \partial^{\mu} \phi + \kappa h'_{\mu\nu} T^{\mu\nu} + \frac{2}{D-2} \kappa \phi T$$

- Scalar couples to the trace of the stress energy tensor and does not decouple. 1.
- Behaves like a Scalar-Tensor/Brans-Dicke Theory. Affects the Newtonian Potential 2.
- vanDam-Veltman-Zakharov Discontinuity. M-> 0 limit not smooth under Stuckleberg 3.

Gravity as an Effective Field Theory : Diffeomorphism and Mass Terms

$$A_{\mu} \to A_{\mu} + \partial_{\mu}\phi$$

5 propagating degrees of freedom 2 transverse + 3 longitudinal

 $\partial_{\nu}T^{\mu\nu}$ 







$$\epsilon_{\pm 2}^{\mu\nu} = \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu},$$

$$\epsilon_{\pm 1}^{\mu\nu} = \frac{1}{\sqrt{2}} [\epsilon_{\pm 1}^{\mu} \epsilon_{0}^{\nu} + \epsilon_{0}^{\mu} \epsilon_{\pm 1}^{\nu}],$$

 $\epsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left[ \epsilon_{+1}^{\mu} \epsilon_{-1}^{\nu} + \epsilon_{-1}^{\mu} \epsilon_{+1}^{\nu} + 2\epsilon_{0}^{\mu} \epsilon_{0}^{\nu} \right]$ 

$$\epsilon^{\mu}_{\pm 1} = \pm \frac{e^{\pm i\phi}}{\sqrt{2}} \left(0, -c_{\theta}c_{\phi} \pm is_{\phi}, -c_{\theta}c_{\phi} \pm is_{\phi}\right)$$

$$\epsilon_0^{\mu} = \frac{E}{m} \left( \sqrt{1 - \frac{m^2}{E^2}}, \hat{p} \right),$$



### $-c_{\theta}s_{\phi} \exists$



Power Counting

- 1. Each external polarization grows as s/m<sup>2</sup>
- 2. Each vertex grows as s
- 3. The propagator grows as 1/s

$$\epsilon_{\pm 2}^{\mu\nu} = \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu},$$

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### $-c_{\theta}s_{\phi} \exists$



Power Counting

- 1. Each external polarization grows as s/m<sup>2</sup>
- 2. Each vertex grows as s
- 3. The propagator grows as 1/s

Amplitude grows as  $\frac{s^3}{m^8 M_{pl}^2}$ Discontinuity as m-> 0. Does not reduce to Einstein-Hilbert action

Unitarity is violated at a scale  $\Lambda_5 = (M_{pl}m^4)^{1/5} \ll M_{pl}$ 

$$\epsilon_{\pm 2}^{\mu\nu} = \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu},$$

$$\epsilon_{\pm 1}^{\mu\nu} = \frac{1}{\sqrt{2}} [\epsilon_{\pm 1}^{\mu} \epsilon_{0}^{\nu} + \epsilon_{0}^{\mu} \epsilon_{\pm 1}^{\nu}],$$

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$$\epsilon_0^{\mu} = \frac{E}{m} \left( \sqrt{1 - \frac{m^2}{E^2}}, \hat{p} \right),$$

De-Rham, Gabadadze, Tolley (2010) Cheung and Remen (2017) Bonifacio, Rosen, Hinterbichler (2019) Georgi, Arkani-Hamed, Schwartz (2001) <u>Schwartz (2003)</u>



# Most general potential

$$S = \frac{1}{2\kappa^2} \int d^D x \left[ \left( \sqrt{-g}R \right) - \sqrt{-g} \frac{1}{4} m^2 V(g,h) \right]$$

Non-linear massive gravity

### Most general potential

$$V_{2}(g,h) = \langle h^{2} \rangle - \langle h \rangle^{2},$$
  

$$V_{3}(g,h) = +c_{1} \langle h^{3} \rangle + c_{2} \langle h^{2} \rangle \langle h \rangle + c_{3} \langle h \rangle^{3},$$
  

$$V_{4}(g,h) = +d_{1} \langle h^{4} \rangle + d_{2} \langle h^{3} \rangle \langle h \rangle + d_{3} \langle h^{2} \rangle^{2} + d_{4} \langle h^{2} \rangle \langle h \rangle^{2} +$$
  

$$V_{5}(g,h) = +f_{1} \langle h^{5} \rangle + f_{2} \langle h^{4} \rangle \langle h \rangle + f_{3} \langle h^{3} \rangle \langle h \rangle^{2} + f_{4} \langle h^{3} \rangle \langle h^{2} \rangle +$$
  

$$+f_{6} \langle h^{2} \rangle \langle h \rangle^{3} + f_{7} \langle h \rangle^{5},$$

Tune coefficients to raise the scale, avoid ghosts

$$S = \frac{1}{2\kappa^2} \int d^D x \left[ \left( \sqrt{-g}R \right) - \sqrt{-g} \frac{1}{4} m^2 V(g,h) \right]$$



### Most general potential

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$$+f_{6} \langle h^{2} \rangle \langle h \rangle^{3} + f_{7} \langle h \rangle^{5},$$

### Tune coefficients to raise the scale, avoid ghosts

$$c_{1} = 2c_{3} + \frac{1}{2}, \qquad c_{2} = -3c_{3} - \frac{1}{2},$$
  

$$d_{1} = -6d_{5} + \frac{3}{2}c_{3} + \frac{5}{16}, \qquad d_{2} = 8d_{5} - \frac{3}{2}c_{3} - \frac{1}{4},$$
  

$$d_{3} = 3d_{5} - \frac{3}{4}c_{3} - \frac{1}{16}, \qquad d_{4} = -6d_{5} + \frac{3}{4}c_{3},$$

De-Rahm, Gabadadze, Tolley Cheung and Remen Bonifiacio, Rosen, Hinterbichler

$$S = \frac{1}{2\kappa^2} \int d^D x \left[ \left( \sqrt{-g}R \right) - \sqrt{-g} \frac{1}{4} m^2 V(g,h) \right]$$



### Most general potential

$$V_{2}(g,h) = \langle h^{2} \rangle - \langle h \rangle^{2},$$
  

$$V_{3}(g,h) = +c_{1} \langle h^{3} \rangle + c_{2} \langle h^{2} \rangle \langle h \rangle + c_{3} \langle h \rangle^{3},$$
  

$$V_{4}(g,h) = +d_{1} \langle h^{4} \rangle + d_{2} \langle h^{3} \rangle \langle h \rangle + d_{3} \langle h^{2} \rangle^{2} + d_{4} \langle h^{2} \rangle \langle h \rangle^{2} +$$
  

$$V_{5}(g,h) = +f_{1} \langle h^{5} \rangle + f_{2} \langle h^{4} \rangle \langle h \rangle + f_{3} \langle h^{3} \rangle \langle h \rangle^{2} + f_{4} \langle h^{3} \rangle \langle h^{2} \rangle +$$
  

$$+f_{6} \langle h^{2} \rangle \langle h \rangle^{3} + f_{7} \langle h \rangle^{5},$$

### Tune coefficients to raise the scale, avoid ghosts

$$c_{1} = 2c_{3} + \frac{1}{2}, \qquad c_{2} = -3c_{3} - \frac{1}{2},$$
  

$$d_{1} = -6d_{5} + \frac{3}{2}c_{3} + \frac{5}{16}, \qquad d_{2} = 8d_{5} - \frac{3}{2}c_{3} - \frac{1}{4},$$
  

$$d_{3} = 3d_{5} - \frac{3}{4}c_{3} - \frac{1}{16}, \qquad d_{4} = -6d_{5} + \frac{3}{4}c_{3},$$

De-Rahm, Gabadadze, Tolley Cheung and Remen Bonifiacio, Rosen, Hinterbichler

# Cut-off scale raised to $~\Lambda_3$ Can show that cut-off can't be raised above $~\Lambda_3$ Realizations of this set up : dRGT gravity, Bi/Multigravity

$$S = \frac{1}{2\kappa^2} \int d^D x \left[ \left( \sqrt{-g}R \right) - \sqrt{-g} \frac{1}{4} m^2 V(g,h) \right]$$





 $\mathcal{S}_{5D}^{EH} \propto \frac{1}{M_5^3} \int d^5 x \sqrt{-g} R_{5D}$ 

High energy growth

Coupled channel analysis



 $s^{3/2}$  /



**A.** Flat Extra dimension compactified on a torus **B.** The Randall Sundrum Model (ADS)

### 5D diffeomorphism with a 5D Planck mass

<u>Compactification (IR phenomenon) should not change the high energy (UV) behavior,</u>





# Understanding the problem : Geometrical Deconstruction of Dimensions

# Understanding the problem : Geometrical Deconstruction of Dimensions



Discretize a dimension with nearest neighbour interaction

$$+ M^2 m^2 \sqrt{g^j} (g^j_{\mu\nu} - g^{j+1}_{\mu\nu}) (g^{\mu\rho}_j g^{\nu\sigma}_j - g^{\mu\nu}_j g^{\rho\sigma}_j) (g^j_{\rho\sigma} - g^{j+1}_{\rho\sigma})$$
ond to broken diffeormorphisms,  $\Lambda_{\min} = (Nm_1^4 M_{\rm Pl})^{1/5}$ 

Formally never recovers the full 5D

# Understanding the problem : Geometrical Deconstruction of Dimensions



truncated KK theory

Discretize a dimension with nearest neighbour interaction

$$+ M^2 m^2 \sqrt{g^j} (g^j_{\mu\nu} - g^{j+1}_{\mu\nu}) (g^{\mu\rho}_j g^{\nu\sigma}_j - g^{\mu\nu}_j g^{\rho\sigma}_j) (g^j_{\rho\sigma} - g^{j+1}_{\rho\sigma})$$
ond to broken diffeormorphisms,  $\Lambda_{\min} = (Nm_1^4 M_{\rm Pl})^{1/5}$ 
of Stuckelberg/Goldstone fields.

Formally never recovers the full 5D

In a full/truncated KK theory, we need all interactions and not just nearest neighbour ones

> Arkani-Hamed, Georgi, Schwartz 2002 Arkani-Hamed, Schwartz 2003 Schwartz 2003





# Compact Extra Dimensions : A primer

e 5-D Ricci scalar expanded as 
$$G^{\mu\nu}R_{MV}$$
 with  $E$  that Dimensions : A primer  
(3.3)  
rel $\Omega$  bifolded1-DiRUS nek mass by  $M_5^3 = M_{Pl}^{2-\alpha}L$ .  
erse and extreminant istogenitive field, outer expansion in poses,  $f_c$  dy  $e^{(i\omega_n y)^*}e^{i\omega_n y} = L\delta_{mn}0 - (1 + \frac{\kappa \delta}{\sqrt{6}})^2)$   
o expand<sup>2</sup> file The define formative field, outer expansion in poses,  $f_c$  dy  $e^{(i\omega_n y)^*}e^{i\omega_n y} = L\delta_{mn}0 - (1 + \frac{\kappa \delta}{\sqrt{6}})^2)$   
 $= \sum_{n=-\infty}^{\infty} h_{\mu\nu,n}(x)e^{i\omega_n y}$   $h_{\mu\nu,n}^* = h_{\mu\nu,-n}, \rho_{\mu,n}^* = \rho_{\mu,-n}, r_n^* = \phi_{-n}$ .  $(1)$   $Z_2 \longrightarrow (3A1)^{-1}$   
 $y) = \sum_{n=-\infty}^{\infty} \rho_{\mu}(x,n)e^{i\omega_n y}$  (3.3)

acterizes the mode number with frequencies  $\omega_n \equiv$ arier expansion imposes,  $\int_0^L dy e^{(i\omega_m y)^*} e^{i\omega_n y} = L\delta_{mn}$ . ensor, vector, and a scalar as  $h_{\mu,\nu}(x)$ ,  $\rho_{\mu}(x)$ , r(x)of the 5-D fields impose,

$$\nu_{,-n}, \ \rho_{\mu,n}^* = \rho_{\mu,-n}, \ r_n^* = \phi_{-n}.$$
(3.4)





y = 0 $y = \pi R$ 

$$(x, y) \stackrel{\text{galto...}}{\longrightarrow} \rho_{\mu}(x, n) e^{i\omega_n y}$$

an integer that characterizes the mode number with frequencies  $\omega_n \equiv$ thogonality of the fourier expansion imposes,  $\int_0^L dy e^{(i\omega_m y)^*} e^{i\omega_n y} = L\delta_{mn}$ . oken to a tensor vector, and scalar as  $h_{\mu\nu}(x)$ ,  $\rho_{\mu}(x)$ , r(x)











y = 0 $y = \pi R$ 

an integer that characterizes the mode number, with frequencies  $\varphi_n \equiv$  $+ h_{\mu\nu,n}^{*} \mathcal{E}^{\mu\nu,\alpha\beta} h_{\alpha\beta,n}^{\prime} - \omega_{n}^{2} \left[ h_{\mu\nu,n}^{\prime} \right]^{2} - \left[ h_{m}^{\prime} \right]_{L}^{2} \right) + \left[ \frac{-1}{4} h_{\mu\nu,n}^{\prime} t_{\mu\nu,n}^{\mu\nu,\nu} + c.c. \right]_{L}^{\prime} h_{\mu\nu,n}^{\prime} + c.c. \right]_{L}^{\prime} h_{\mu\nu,n}^{\prime} = I \delta_{mn}.$ 

> tensor vector and scalar as h  $(r) \quad o \quad (r)$










an integer that characterizes the mode number, with frequencies  $\varphi_n \equiv$  $+ h_{\mu\nu,n}^{*} \mathcal{E}^{\mu\nu,\alpha\beta} h_{\alpha\beta,n}^{\prime} - \omega_{n}^{2} \left[ h_{\mu\nu,n}^{\prime} \right]^{2} - \left[ h_{m}^{\prime} \right]_{L}^{2} \right) + \left[ \frac{-1}{4} h_{\mu\nu,n}^{\prime} t_{\mu\nu,n}^{\mu\nu,\nu} + c.c. \right]_{L}^{\prime} h_{\mu\nu,n}^{\prime} + c.c. \right]_{L}^{\prime} h_{\mu\nu,n}^{\prime} = I \delta_{mn}.$ 

> tensor vector and scalar as h  $(r) \quad o \quad (r)$







![](_page_37_Picture_3.jpeg)

![](_page_37_Picture_4.jpeg)

an integer that characterizes the mode number, with frequencies  $\varphi_n \equiv$  $+ \sum_{\mu\nu,n} h_{\mu\nu,n}^{*} \mathcal{E}^{\mu\nu,\alpha\beta} h_{\alpha\beta,n}^{\prime} - \omega_{n}^{\prime} (h_{\mu\nu,n}^{\prime})^{2} - |h_{m}^{\prime}|_{L}^{2} ) + \left[ \frac{-1}{44} h_{\mu\nu,n}^{\prime} h_{\mu\nu,n}^{\prime} + c.c. \right]_{l}^{\prime} h_{\mu\nu,n}^{\prime} + c.c. \right]_{l}^{\prime} h_{\mu\nu,n}^{\prime} = I \int_{mn}^{l} h_{\mu\nu,n}^{\prime} h_{\mu\nu,n}^{\prime} + c.c. \int_{mn}^{l} h_{\mu\nu,n}^{\prime} h_{\mu\nu,n}^{\prime} h_{\mu\nu,n}^{\prime} + c.c. \int_{mn}^{l} h_{\mu\nu,n}^{\prime} h$ 

> tensor rector and scalar as h  $r(x), o_{x}(x), r(x)$

![](_page_37_Figure_10.jpeg)

![](_page_37_Figure_11.jpeg)

# Compact Extra Dimensions : A Primer

# Compact Extra Dimensions : A Primer

- A 5D field appears as a tower of KK modes from 4D point of view, with each mode having a profile (i) in the extra dimension.
- The profiles and the KK masses are obtained by solving an eigenvalue problem (or wave equations in (ii) 5D space-time).
- The coupling of particles (i.e., zero and KK modes) is proportional to the overlap of their profiles in (iii) the extra dimension.

![](_page_39_Figure_7.jpeg)

# Compact Extra Dimensions : A Primer

- A 5D field appears as a tower of KK modes from 4D point of view, with each mode having a profile (i) in the extra dimension.
- The profiles and the KK masses are obtained by solving an eigenvalue problem (or wave equations in (ii)5D space-time).
- The coupling of particles (i.e., zero and KK modes) is proportional to the overlap of their profiles in (iii)the extra dimension.
  - The 4D graviton and its KK modes
  - 4D vectors and their KK modes

![](_page_40_Figure_6.jpeg)

![](_page_40_Figure_7.jpeg)

![](_page_40_Figure_8.jpeg)

Vanishes on Orbifold, Higher Modes get Higgsed

Higher modes of scalar get Higgsed, leaving behind a single field with a flat wave function (radion/dilaton/modulus/breathing mode)

![](_page_40_Figure_14.jpeg)

# Compact Extra Dimensions : Randall-S

![](_page_41_Figure_1.jpeg)

![](_page_42_Figure_1.jpeg)

![](_page_42_Picture_2.jpeg)

L

![](_page_43_Figure_1.jpeg)

![](_page_43_Figure_3.jpeg)

Mass spectrum

$$m_n = k x_n e^{-k r_c \pi} \quad \text{TeV}$$

scale masses for  $kr_c = 11 - 12$ 

$$S = \int d^4 x [dy \mathcal{L}_{5D}] \equiv \int d^4 x \mathcal{L}_{4D}^{(\text{eff})}$$

$$\hat{h}_{\mu\nu}(x,y) = \frac{1}{\sqrt{\pi r_c}} \sum_{n=0}^{+\infty} \hat{h}_{\mu\nu}^{(n)}(x) \psi_n(\varphi)$$

$$\left|\underbrace{\hat{r}(x)}_{\text{5D field}} = \frac{1}{\sqrt{\pi r_c}} \underbrace{\hat{r}^{(0)}(x)}_{\text{4D fields wfxn}} \underbrace{\psi_0}_{\text{wfxn}}\right|$$

# Radion KK Decomposition

![](_page_45_Picture_5.jpeg)

\_

$$S = \int d^4x [dy \mathcal{L}_{5D}] \equiv \int d^4x \mathcal{L}_{4D}^{(\text{eff})}$$

$$\hat{h}_{\mu\nu}(x,y) = \frac{1}{\sqrt{\pi r_c}} \sum_{n=0}^{+\infty} \hat{h}_{\mu\nu}^{(n)}(x) \psi_n(\varphi)$$

- is orthonormal+complete with discrete spectrum  $\mu_n$
- takes 5D graviton  $\rightarrow$  4D graviton + massive spin-2 tower

Describes a Hilbert space of wave Functions

$$\underbrace{\hat{r}(x)}_{\text{5D field}} = \frac{1}{\sqrt{\pi r_c}} \underbrace{\hat{r}^{(0)}(x)}_{\text{4D fields wfxn}} \underbrace{\psi_0}_{\text{wfxn}}$$

## Radion KK Decomposition

![](_page_46_Picture_9.jpeg)

$$S = \int d^4x [dy \mathcal{L}_{5D}] \equiv \int d^4x \mathcal{L}_{4D}^{(\text{eff})}$$

$$\hat{h}_{\mu\nu}(x,y) = \frac{1}{\sqrt{\pi r_c}} \sum_{n=0}^{+\infty} \hat{h}_{\mu\nu}^{(n)}(x) \psi_n(\varphi)$$

- is orthonormal+complete with discrete spectrum  $\mu_n$
- takes 5D graviton  $\rightarrow$  4D graviton + massive spin-2 tower

Describes a Hilbert space of wave Functions

# Completeness Schrodinger Equation Orthonormality $-\frac{d}{dy}\left[e^{-4k|y|}\frac{d\psi_{n}}{dy}\right] = m_{n}^{2}e^{-2k|y|}\psi_{n} \qquad \left|\frac{1}{\pi r_{c}}\int_{-\pi r_{c}}^{+\pi r_{e}}dye^{-2k|y|}\psi_{m}(y)\psi_{n}(y) = \delta_{mn}\right| \qquad \left|\frac{1}{\pi r_{c}}e^{-2k|y|}\sum_{j}\psi_{j}(y)\psi_{j}(y') = \delta(y-y')\right|$

$$\left| \underbrace{\hat{r}(x)}_{\text{5D field}} = \frac{1}{\sqrt{\pi r_c}} \underbrace{\hat{r}^{(0)}(x)}_{\text{4D fields wfxn}} \underbrace{\psi_0}_{\text{wfxn}} \right|$$

![](_page_47_Picture_9.jpeg)

$$S = \int d^4x [dy \mathcal{L}_{5D}] \equiv \int d^4x \mathcal{L}_{4D}^{(\text{eff})}$$

$$\hat{h}_{\mu\nu}(x,y) = \frac{1}{\sqrt{\pi r_c}} \sum_{n=0}^{+\infty} \hat{h}_{\mu\nu}^{(n)}(x) \psi_n(\varphi)$$

- is orthonormal+complete with discrete spectrum  $\mu_n$
- takes 5D graviton  $\rightarrow$  4D graviton + massive spin-2 tower
- Describes a Hilbert space of wave Functions

![](_page_48_Figure_6.jpeg)

Torus 
$$\psi_n = \begin{cases} \psi_0 = \frac{1}{\sqrt{2}} \\ \psi_n = -\cos(n|\varphi|) \end{cases}$$
 ADS

$$\mu_n = m_n r_c = n$$

$$\left| \underbrace{\hat{r}(x)}_{\text{5D field}} = \frac{1}{\sqrt{\pi r_c}} \underbrace{\hat{r}^{(0)}(x)}_{\text{4D fields wfxn}} \underbrace{\psi_0}_{\text{wfxn}} \right|$$

# Completeness $-\frac{d}{dy}\left[e^{-4k|y|}\frac{d\psi_{n}}{dy}\right] = m_{n}^{2}e^{-2k|y|}\psi_{n} \qquad \left|\frac{1}{\pi r_{c}}\int_{-\pi r_{c}}^{+\pi r_{e}}dy e^{-2k|y|}\psi_{m}(y)\psi_{n}(y) = \delta_{mn}\right| \qquad \left|\frac{1}{\pi r_{c}}e^{-2k|y|}\sum_{j}\psi_{j}(y)\psi_{j}(y') = \delta(y-y')\right|$

$$\psi_{0} = \sqrt{\frac{\pi k r_{c}}{1 - e^{-2\pi k r_{c}}}}$$
$$\psi_{n}(y) = \frac{e^{+2k|y|}}{N_{n}} \left[ J_{2} \left( \frac{m_{n}}{k} e^{+k|y|} \right) + b_{n2} Y_{2} \left( \frac{m_{n}}{k} e^{+k|y|} \right) \right]$$

$$m_n = k x_n e^{-kr_c \pi}$$

![](_page_48_Picture_15.jpeg)

### An Elastic scattering process in compactified theories

![](_page_50_Figure_2.jpeg)

![](_page_50_Figure_4.jpeg)

$$R_{5\mathrm{D}} = \tilde{G}^{MN} R_{MN}$$

### An Elastic scattering process in compactified theories

![](_page_51_Figure_2.jpeg)

![](_page_51_Figure_3.jpeg)

$$R_{5D} = \tilde{G}^{MN} R_{MN}$$
Expansion of Ricci gives two different  
types of coupling structures
$$n_{3}$$

$$n_{4}$$

$$n_{1} \rightarrow a_{n_{1}n_{2}n_{3}n_{4}} \quad b_{\mathcal{P}[n_{1},n_{2},n_{3},n_{4}]}$$

$$n_{1} \rightarrow c \rightarrow b_{n_{1}n_{2}r}$$

### An Elastic scattering process in compactified theories

![](_page_52_Figure_2.jpeg)

![](_page_52_Figure_3.jpeg)

Define  $\mathcal{L}_{h^H r^R}^{(\text{RS})} \equiv \text{all terms in } \mathcal{L}_{5D}^{(\text{RS})}$  with *H* graviton fields and *R* radion fields. By construction, each term in this set is either

- A-Type: has two spatial derivatives  $\partial_{\mu}\partial_{\nu}$ , or
- **B-Type:** has two extra-dimensional derivatives  $\partial_u^2$

$$\begin{aligned} \mathcal{L}_{h^{H}r^{R}}^{(\mathrm{RS})} &= \mathcal{L}_{A:h^{H}r^{R}}^{(\mathrm{RS})} + \mathcal{L}_{B:h^{H}r^{R}}^{(\mathrm{RS})} \\ &= \kappa^{(H+R-2)} \left[ e^{k[2(R-1)|y| - R\pi r_{c}]} \ \overline{\mathcal{L}}_{A:h^{H}r^{R}}^{(\mathrm{RS})} + e^{i t_{A}} \right] \end{aligned}$$

$$R_{5D} = \tilde{G}^{MN} R_{MN}$$
Expansion of Ricci gives two different  
types of coupling structures
$$n_{3}$$

$$\sum a_{n_{1}n_{2}n_{3}n_{4}} b_{\mathcal{P}[n_{1},n_{2},n_{3},n_{4}]} \qquad \boxed{n_{1}}_{n_{2}} \stackrel{r}{\sim} \sum b_{n_{1}n_{2}r}$$

 $e^{k[2(R-2)|y|-R\pi r_c]} \overline{\mathcal{L}}_{B:h^H r^R}^{(RS)}$ 

![](_page_54_Figure_1.jpeg)

![](_page_55_Figure_1.jpeg)

Couplings are proportional to products/overlaps of wave functions and derivatives of them

![](_page_56_Figure_1.jpeg)

Couplings are proportional to products/overlaps of wave functions and derivatives of them

![](_page_56_Figure_3.jpeg)

$$a_{nnj} = \frac{1}{\pi r_c} \int_{-\pi r_c}^{\pi r_c} dy e^{-2k|y|} \psi_n(y) \psi_n(y) \psi_j(y)$$

$$b_{nnj} = \frac{r_c}{\pi} \int_{-\pi r_c}^{\pi r_c} dy e^{-4k|y|} (\partial_y \psi_n(y) \partial_y \psi_n(y)) \psi_j(y)$$

![](_page_57_Figure_1.jpeg)

Couplings are proportional to products/overlaps of wave functions and derivatives of them

![](_page_57_Figure_3.jpeg)

![](_page_57_Figure_4.jpeg)

3 and 4 point non-derivative couplings

$$a_{nnj} = \frac{1}{\pi r_c} \int_{-\pi r_c}^{\pi r_c} dy e^{-2k|y|} \psi_n(y) \psi_n(y) \psi_j(y)$$

$$a_{nnnn} = \frac{1}{\pi r_c} \int_{-\pi r_c}^{\pi r_c} dy e^{-2k|y|}$$

3 and 4 point derivative couplings

$$b_{nnj} = \frac{r_c}{\pi} \int_{-\pi r_c}^{\pi r_c} dy e^{-4k|y|} (\partial_y \psi_n(y) \partial_y \psi_n(y)) \psi_j(y)$$

$$b_{nnnn} = \frac{r_c}{\pi} \int_{-\pi r_c}^{\pi r_c} dy e^{-4k|y|} (\partial_y \psi_n(y) \partial_y \psi_n(y)) \psi_n(y) \psi_n(y)$$

 $a_{n_1n_2n_3n_4}$   $b_{\mathcal{P}[n_1,n_2,n_3,n_4]}$ 

$$n_1 \rightarrow r \quad \supset \quad b$$
$$n_2 \rightarrow r \quad b$$

 $|\psi_n(y)\psi_n(y)\psi_n(y)\psi_n(y)|$ 

![](_page_57_Picture_14.jpeg)

![](_page_58_Figure_1.jpeg)

Couplings are proportional to products/overlaps of wave functions and derivatives of them

![](_page_58_Figure_3.jpeg)

![](_page_58_Figure_4.jpeg)

3 and 4 point non-derivative couplings

$$a_{nnj} = \frac{1}{\pi r_c} \int_{-\pi r_c}^{\pi r_c} dy e^{-2k|y|} \psi_n(y) \psi_n(y) \psi_j(y)$$

$$a_{nnnn} = \frac{1}{\pi r_c} \int_{-\pi r_c}^{\pi r_c} dy e^{-2k|y|}$$

### 3 and 4 point derivative couplings

$$b_{nnj} = \frac{r_c}{\pi} \int_{-\pi r_c}^{\pi r_c} dy e^{-4k|y|} (\partial_y \psi_n(y) \partial_y \psi_n(y)) \psi_j(y)$$

$$b_{nnnn} = \frac{r_c}{\pi} \int_{-\pi r_c}^{\pi r_c} dy e^{-4k|y|} (\partial_y \psi_n(y) \partial_y \psi_n(y)) \psi_n(y) \psi_n(y)$$

$$b_{nnr} = \frac{r_c}{\pi} e^{-\pi k r_c} \int_{-\pi r_c}^{\pi r_c} dy e^{-2k|y|} (\partial_y \psi_n(y) \partial_y \psi_n(y)) \psi_0$$

 $a_{n_1n_2n_3n_4}$   $b_{\mathcal{P}[n_1,n_2,n_3,n_4]}$ 

$$n_1 \rightarrow r \quad \supset \quad b$$
$$n_2 \rightarrow r \quad b$$

 $|\psi_n(y)\psi_n(y)\psi_n(y)\psi_n(y)|$ 

![](_page_58_Picture_15.jpeg)

![](_page_59_Figure_1.jpeg)

![](_page_60_Figure_1.jpeg)

 $\frac{1}{\pi r_c} e^{-2k|y|} \sum_{i} \psi_i(y) \psi_i(y') = \delta(y - y')$ 

![](_page_60_Figure_5.jpeg)

![](_page_61_Figure_1.jpeg)

Product of two 3 point a couplings is a 4 point a-type coupling

$$\sum_{j} a_{nnj}^2 \quad a_{nnj} = \frac{1}{\pi r_c} \int_{-\pi r_c}^{\pi r_c} dy e^{-2k|y|} \psi_n(y) \psi_n(y)$$

$$\frac{1}{\pi r_c} e^{-2k|y|} \sum_j \psi_j(y) \psi_j(y') = \delta(y - y')$$

![](_page_61_Figure_6.jpeg)

![](_page_61_Figure_7.jpeg)

![](_page_62_Figure_1.jpeg)

Product of two 3 point a couplings is a 4 point a-type coupling

$$m_n^2 a_{nnnn} = \frac{3}{4} \sum_j m_j^2 a_{nnj}^2 \qquad b_{nnnn} = \frac{1}{3} m_n^2 r_c^2 a_{nnnn}$$

$$\begin{bmatrix} \sum_{j} a_{nnj}^{2} & a_{nnj} = \frac{1}{\pi r_{c}} \int_{-\pi r_{c}}^{\pi r_{c}} dy e^{-2k|y|} \psi_{n}(y) \psi_{n}(y) \\ \frac{1}{\pi r_{c}} e^{-2k|y|} \sum_{j} \psi_{j}(y) \psi_{j}(y') = \delta(y - y') \\ \begin{bmatrix} a_{nnnn} = \sum_{j} a_{nnj}^{2} \end{bmatrix}$$

![](_page_62_Figure_5.jpeg)

![](_page_64_Figure_1.jpeg)

![](_page_65_Figure_1.jpeg)

$$\mathcal{M}(s,\cos\theta) \equiv \sum_{k=-\infty}^{+5} \overline{\mathcal{M}}^{(k)}(\cos\theta) \cdot s^k$$

$$\epsilon_{\pm 2}^{\mu\nu} = \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu},$$

$$\epsilon_{\pm 1}^{\mu\nu} = \frac{1}{\sqrt{2}} \left[ \epsilon_{\pm 1}^{\mu} \epsilon_{0}^{\nu} + \epsilon_{0}^{\mu} \epsilon_{\pm 1}^{\nu} \right],$$

$$\epsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left[ \epsilon_{+1}^{\mu} \epsilon_{-1}^{\nu} + \epsilon_{-1}^{\mu} \epsilon_{+1}^{\nu} + \right]$$

$$\epsilon_{\pm 1}^{\mu} = \pm \frac{e^{\pm i\phi}}{\sqrt{2}} \left(0, -c_{\theta}c_{\phi} \pm is\right)$$

$$\epsilon_0^{\mu} = \frac{E}{m} \left( \sqrt{1 - \frac{m^2}{E^2}}, \hat{p} \right),$$

![](_page_65_Picture_9.jpeg)

![](_page_66_Figure_1.jpeg)

$$\mathcal{M}(s,\cos\theta) \equiv \sum_{k=-\infty}^{+5} \overline{\mathcal{M}}^{(k)}(\cos\theta) \cdot s^k$$

$$m_n^2 a_{nnnn} = \frac{3}{4} \sum_j m_j^2 a_{nnj}^2$$

$$\epsilon_{\pm 2}^{\mu\nu} = \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu},$$

$$\epsilon_{\pm 1}^{\mu\nu} = \frac{1}{\sqrt{2}} [\epsilon_{\pm 1}^{\mu} \epsilon_{0}^{\nu} + \epsilon_{0}^{\mu} \epsilon_{\pm 1}^{\nu}],$$

$$\frac{1}{\alpha_{nnj}^{2}} \qquad \epsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left[ \epsilon_{\pm 1}^{\mu} \epsilon_{-1}^{\nu} + \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu} + 2\epsilon_{0}^{\mu} \epsilon_{0}^{\nu} \right] \\
\epsilon_{\pm 1}^{\mu} = \pm \frac{e^{\pm i\phi}}{\sqrt{2}} \left( 0, -c_{\theta}c_{\phi} \pm is_{\phi}, -c_{\theta}s_{\phi} + is_{\phi} \right) + \epsilon_{\pm 1}^{\mu\nu} \left[ 0, -c_{\theta}c_{\phi} \pm is_{\phi} \right] \\$$

$$\epsilon_0^{\mu} = \frac{E}{m} \left( \sqrt{1 - \frac{m^2}{E^2}}, \hat{p} \right),$$

![](_page_66_Picture_11.jpeg)

![](_page_67_Figure_1.jpeg)

$$\mathcal{M}^{(4)}(\cos\theta) = \frac{\kappa^2}{\pi r_c} \frac{(7+\cos 2\theta)^2}{27648m_n^8} \cdot \left(4m_n^2 a_{nnnn} - 3\sum_j m_j^2 a_n^2\right)$$
$$\mathcal{M}^{(4)}(\cos\theta) = \frac{\kappa^2}{\pi r_c} \frac{\sin^2\theta}{3456m_n^8} \cdot \left(-108\frac{b_{nnr}^2}{r_c^4} + 12m_n^4 a_{nn0}^2 - 16m_n^4 a_{nnnn} + 15\sum_j m_j^4 a_n^2\right)$$

$$\mathcal{M}(s,\cos\theta) \equiv \sum_{k=-\infty}^{+5} \overline{\mathcal{M}}^{(k)}(\cos\theta) \cdot s^k$$

$$m_n^2 a_{nnnn} = \frac{3}{4} \sum_j m_j^2 a_{nnj}^2$$

$$\epsilon_{\pm 2}^{\mu\nu} = \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu},$$

$$\epsilon_{\pm 1}^{\mu\nu} = \frac{1}{\sqrt{2}} [\epsilon_{\pm 1}^{\mu} \epsilon_{0}^{\nu} + \epsilon_{0}^{\mu} \epsilon_{\pm 1}^{\nu}],$$

![](_page_67_Figure_8.jpeg)

$$\epsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left[ \epsilon_{+1}^{\mu} \epsilon_{-1}^{\nu} + \epsilon_{-1}^{\mu} \epsilon_{+1}^{\nu} + \frac{i\phi}{\sqrt{6}} \right]$$

$$\epsilon_{\pm 1}^{\mu} = \pm \frac{e^{\pm i\varphi}}{\sqrt{2}} \left(0, -c_{\theta}c_{\phi} \pm is\right)$$

$$\epsilon_0^{\mu} = \frac{E}{m} \left( \sqrt{1 - \frac{m^2}{E^2}}, \hat{p} \right),$$

![](_page_67_Picture_12.jpeg)

![](_page_68_Figure_1.jpeg)

$$\mathcal{M}^{(4)}(\cos\theta) = \frac{\kappa^2}{\pi r_c} \frac{(7+\cos 2\theta)^2}{27648m_n^8} \cdot \left(4m_n^2 a_{nnnn} - 3\sum_j m_j^2 a_n^2\right)$$
$$\mathcal{M}^{(4)}(\cos\theta) = \frac{\kappa^2}{\pi r_c} \frac{\sin^2\theta}{3456m_n^8} \cdot \left(-108\frac{b_{nnr}^2}{r_c^4} + 12m_n^4 a_{nn0}^2 - 16m_n^4 a_{nnnn} + 15\sum_j m_j^4 a_n^2\right)$$
$$\mathcal{O}(S^2) \qquad \mathcal{M}^{(2)}(\cos\theta) = -\frac{\kappa^2}{\pi r_c} \frac{(7+\cos 2\theta)}{5184m_n^8} \cdot \left(-108\frac{b_{nnr}^2}{r_c^4} + 12m_n^4 a_{nn0}^2 - 16m_n^4 a_{nnnn} + 15\sum_j m_j^4 a_n^2\right)$$

$$\mathcal{M}(s,\cos\theta) \equiv \sum_{k=-\infty}^{+5} \overline{\mathcal{M}}^{(k)}(\cos\theta) \cdot s^k$$

$$m_n^2 a_{nnnn} = \frac{3}{4} \sum_j m_j^2 a_{nnj}^2$$

$$\epsilon_{\pm 2}^{\mu\nu} = \epsilon_{\pm 1}^{\mu} \epsilon_{\pm 1}^{\nu},$$

$$\epsilon_{\pm 1}^{\mu\nu} = \frac{1}{\sqrt{2}} \left[ \epsilon_{\pm 1}^{\mu} \epsilon_{0}^{\nu} + \epsilon_{0}^{\mu} \epsilon_{\pm 1}^{\nu} \right],$$

![](_page_68_Figure_8.jpeg)

$$m_j^6 a_{nnj}^2 = 5 m_n^2 \sum_j m_j^4 a_{nnj}^2 - rac{16}{3} m_n^6 a_{nnnn}$$

$$\epsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left[ \epsilon_{+1}^{\mu} \epsilon_{-1}^{\nu} + \epsilon_{-1}^{\mu} \epsilon_{+1}^{\nu} + \frac{1}{\sqrt{6}} \right]$$

$$\epsilon_{\pm 1}^{\mu} = \pm \frac{e^{\pm i\phi}}{\sqrt{2}} \left(0, -c_{\theta}c_{\phi} \pm is\right)$$

$$\epsilon_0^{\mu} = \frac{E}{m} \left( \sqrt{1 - \frac{m^2}{E^2}}, \hat{p} \right),$$

![](_page_68_Picture_13.jpeg)

Elastic 2-to-2 KK Mode Scattering Matrix Elements in RS1 Fastest Energy Growth per Helicity Combination:  $(\lambda_1, \lambda_2) \rightarrow (\lambda_3, \lambda_4)$ 

				λ			**BEFORE SUM RULES
		-2	-1	0	ቍ1	+2	
	-2	2343234543456543454323432	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	2343234543456543454323432	
	-1	3454345654567654565434543	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\lambda_4$ -2 -1 0 +1
λ	0	4565456765678765676545654	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	67876789878910987898767876	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $
	<b>÷1</b>	3454345654567654565434543	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	5       6       7       6       5         6       7       8       7       6         7       8       9       8       7         6       7       8       7       6         5       6       7       8       5	4       5       6       5       4         5       6       7       6       5         6       7       8       7       6         5       6       7       6       5         4       5       6       5       4	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	+2 <mark>6787</mark>
	+2	2       3       4       3       2         3       4       5       4       3         4       5       6       5       4         3       4       5       4       3         2       3       4       5       4       3	3       4       5       4       3         4       5       6       5       4         5       6       7       6       5         4       5       6       5       4         3       4       5       4       3	4       5       6       5       4         5       6       7       6       5         6       7       8       7       6         5       6       7       6       5         4       5       6       5       4	3       4       5       4       3         4       5       6       5       4         5       6       7       6       5         4       5       6       5       4         3       4       5       6       5       4	2 3 4 3 2 3 4 5 4 3 4 5 6 5 4 3 4 5 4 3 2 3 4 5 2	Legend X X = O(E <sup>x</sup> ) growt

![](_page_70_Picture_3.jpeg)

![](_page_70_Picture_4.jpeg)

![](_page_70_Picture_5.jpeg)

Elastic 2-to-2 KK Mode Scattering Matrix Elements in RS1

![](_page_71_Figure_2.jpeg)

Phys. Rev. D 107, (2023) 03505, Phys. Rev. D 101 (2020) 7, 075013 Phys.Rev.D 101 (2020) 5, 055013 Phys. Rev. D 103 (2022), 095024 Phys. Rev. D 107, (2023) 03505,

Can be extended to a Full Goldberger-Wise stabilized Model. 1.

Matter in Bulk or brane scatterings with KK gravitons get unitarized by a similar mechanisms 2.

![](_page_71_Picture_7.jpeg)

![](_page_71_Figure_9.jpeg)
$$G_{MN} = e^{2A(z)} \begin{pmatrix} e^{-\kappa\varphi/\sqrt{6}}(\eta_{\mu\nu} + \kappa h_{\mu\nu}) & \frac{\kappa}{\sqrt{2}}A_{\mu} \\ \frac{\kappa}{\sqrt{2}}A_{\mu} & -\left(1 + \frac{\kappa}{\sqrt{6}}\varphi\right)^2 \end{pmatrix} dz \\ \hline h_{\mu\nu}(x^{\alpha}, z) = \sum_{n=0}^{\infty} h_{\mu\nu}^{(n)}(x^{\alpha})f^{(n)}(z) \\ A_{\mu}(x^{\alpha}, z) = \sum_{n=1}^{\infty} A_{\mu}^{(n)}(x^{\alpha})g^{(n)}(z) \\ \varphi(x^{\alpha}, z) = \sum_{n=0}^{\infty} \varphi^{(n)}(x)k^{(n)}(z) \\ \end{pmatrix}$$

 $ds^{2} = e^{2A(z)} (\eta_{\mu\nu} dx^{\mu} dx^{\nu} - dz^{2})$ 



$$G_{MN} = e^{2A(z)} \begin{pmatrix} e^{-\kappa\varphi/\sqrt{6}}(\eta_{\mu\nu} + \kappa h_{\mu\nu}) & \frac{\kappa}{\sqrt{2}}A_{\mu} \\ \frac{\kappa}{\sqrt{2}}A_{\mu} & -\left(1 + \frac{\kappa}{\sqrt{6}}\varphi\right)^{2} \end{pmatrix} \begin{bmatrix} d \\ d \\ h_{\mu\nu}(x^{\alpha}, z) = \sum_{n=0}^{\infty} h_{\mu\nu}^{(n)}(x^{\alpha})f^{(n)}(z) \\ A_{\mu}(x^{\alpha}, z) = \sum_{n=1}^{\infty} A_{\mu}^{(n)}(x^{\alpha})g^{(n)}(z) \\ \varphi(x^{\alpha}, z) = \sum_{n=0}^{\infty} \varphi^{(n)}(x)k^{(n)}(z) \end{bmatrix} \xrightarrow{\bullet} \text{Only 0 mode}$$

 $ds^{2} = e^{2A(z)} (\eta_{\mu\nu} dx^{\mu} dx^{\nu} - dz^{2})$ 



- K gravitons, O mode + tower of massive states
- I KK graviphotons/goldstones
- e/radion is physical, higher modes higgsed/goldstones

$$G_{MN} = e^{2A(z)} \begin{pmatrix} e^{-\kappa\varphi/\sqrt{6}}(\eta_{\mu\nu} + \kappa h_{\mu\nu}) & \frac{\kappa}{\sqrt{2}}A_{\mu} \\ \frac{\kappa}{\sqrt{2}}A_{\mu} & -\left(1 + \frac{\kappa}{\sqrt{6}}\varphi\right)^2 \end{pmatrix} \begin{bmatrix} d \\ d \\ h_{\mu\nu}(x^{\alpha}, z) = \sum_{n=0}^{\infty} h_{\mu\nu}^{(n)}(x^{\alpha})f^{(n)}(z) \\ A_{\mu}(x^{\alpha}, z) = \sum_{n=1}^{\infty} A_{\mu}^{(n)}(x^{\alpha})g^{(n)}(z) \\ \varphi(x^{\alpha}, z) = \sum_{n=0}^{\infty} \varphi^{(n)}(x)k^{(n)}(z) \end{bmatrix} \xrightarrow{} Only 0 \text{ mode}$$

$$\mathcal{L} = \mathcal{L}_{EH} + \mathcal{L}_{CC} + \Delta \mathcal{L} + \mathcal{L}_{GF} + \mathcal{L}_{m}$$

$$\mathcal{L}_2 = \frac{1}{2} h^n_{\mu\nu} \mathcal{D}^{\mu\nu\rho\sigma}_h h^n_{\rho\sigma} + \frac{1}{2} A^n_\mu \mathcal{D}^{\mu\nu}_A A^n_\nu + \frac{1}{2} \varphi D_\varphi \varphi$$

$$\mathcal{L}_{\rm GF} = F^{\mu}F_{\mu} - F_5F_5,$$

$$F_{\mu}^n = -\left(\partial^{\nu}h_{\mu\nu}^n - \frac{1}{2}\partial_{\mu}h^n + \frac{1}{\sqrt{2}}m_nA_{\mu}^n\right),$$

$$F_5^n = -\left(\frac{1}{2}m_nh^n - \frac{1}{\sqrt{2}}\partial^{\mu}A_{\mu}^n + \sqrt{\frac{3}{2}}m_n\varphi^n\right)$$

 $ds^{2} = e^{2A(z)} (\eta_{\mu\nu} dx^{\mu} dx^{\nu} - dz^{2})$ 

$$A(z) = -\ln(kz)$$

- K gravitons, O mode + tower of massive states
- KK graviphotons/goldstones
- e/radion is physical, higher modes higgsed/goldstones

Gauge Fixing Lagrangian



Ward Identities 
$$\begin{array}{l} \langle \mathbf{T}F_{\mu}(x)\Phi\rangle = \langle \mathbf{T}F_{5}(x)\Phi\rangle = 0 \\ \\ \langle \mathbf{T}\left(\partial^{\nu}(h_{\mu\nu}^{n} - \frac{1}{2}\eta_{\mu\nu}h^{n}) + \frac{1}{\sqrt{2}}m_{n}A_{\mu}^{n}\right) \\ \\ \langle \mathbf{T}\left(\frac{1}{2}m_{n}h^{n} - \frac{1}{\sqrt{2}}\partial^{\mu}A_{\mu}^{n} + \sqrt{\frac{3}{2}}m_{n}\varphi^{n}\right) \end{array}$$

Any External on-shell physical fields after LSZ amputation



Ward Identities 
$$\begin{array}{l} \langle \mathbf{T}F_{\mu}(x)\Phi\rangle = \langle \mathbf{T}F_{5}(x)\Phi\rangle = 0 \\ \\ \langle \mathbf{T}\left(\partial^{\nu}(h_{\mu\nu}^{n} - \frac{1}{2}\eta_{\mu\nu}h^{n}) + \frac{1}{\sqrt{2}}m_{n}A_{\mu}^{n}\right) \\ \\ \langle \mathbf{T}\left(\frac{1}{2}m_{n}h^{n} - \frac{1}{\sqrt{2}}\partial^{\mu}A_{\mu}^{n} + \sqrt{\frac{3}{2}}m_{n}\varphi^{n}\right) \end{array}$$

Any External on-shell physical fields after LSZ amputation











Consider the longitudinal polarizations of the KK graviton

$$\begin{split} \varepsilon_{0}^{\mu\nu} &= \frac{1}{\sqrt{6}} \left( \varepsilon_{+}^{\mu} \varepsilon_{-}^{\nu} + \varepsilon_{-}^{\mu} \varepsilon_{+}^{\nu} + 2\varepsilon_{0}^{\mu} \varepsilon_{0}^{\nu} \right) & \text{The bad high end} \\ \text{Jsing the polarization sum} & \sum_{\lambda=\pm,0} \varepsilon_{\lambda}^{\mu} \varepsilon_{\lambda}^{\nu*} = -\eta_{\mu\nu} + \frac{p^{\mu}}{m} \\ \text{Rewrite the longitudinal polarizations as} & \varepsilon_{0}^{\mu\nu} = \frac{1}{\sqrt{6}} \left( \\ \text{And separate out the bad high energy behaviour} & \varepsilon_{0}^{\mu} \equiv \frac{p}{n} \\ & \downarrow \\ \text{Bad High Energy g} \\ \varepsilon_{0}^{\mu\nu} &= \tilde{\varepsilon}_{0}^{\mu\nu} + \frac{1}{\sqrt{6}} \left( \eta^{\mu\nu} + 2 \frac{p^{\mu}p^{\nu}}{m^{2}} + 3 \frac{p^{\mu}\tilde{\varepsilon}_{0}^{\nu} + p^{\nu}\tilde{\varepsilon}_{0}^{\mu}}{m} \right) \end{split}$$

 $E^2$ 

energy growth part comes from the last term  $~~\epsilon_0^{\mu
u}\sim {\cal O}(E^2/m^2)$  .





Use Ward Identities to write the scattering amplitude as

$$T^{h}_{\mu\nu}\epsilon^{\mu\nu}_{0} = T^{\varphi} + \mathcal{O}(s^{0})$$

KK Goldstone Theorem :

Scattering Amplitude of longitudinally polarized KK graviton=Scalar Goldstones in the High Energy Limit

$$T^h_{\mu\nu}\epsilon^{\mu\nu}_{\pm 1} = -iT^A_\mu\epsilon^\mu_\pm + \mathcal{O}(s^0)$$

 $\mathcal{O}(m^2/E^2)$ 

Similarly for Helicity 1 one states

$$\begin{split} & \underbrace{\varphi_{(n_2)}^{(n_2)} \qquad \varphi_{(n_3)}^{(n_4)} \qquad \sum_{S^{(n_3)}}^{(n_3)} \qquad \varphi_{(n_4)}^{(n_4)} \qquad \sum_{S^{(n_2)}}^{(n_2)} \qquad \varphi_{(n_4)}^{(n_4)} \qquad \sum_{S^{(n_2)}}^{(n_2)} \qquad \varphi_{(n_4)}^{(n_4)} \qquad \sum_{S^{(n_2)}}^{(n_3)} \qquad \varphi_{(n_4)}^{(n_4)} \qquad \sum_{S^{(n_2)}}^{(n_3)} \qquad \sum_{P_2 \leq P_1}^{(n_3)} \qquad \sum_{P_3 \leq P_1}^{(n_3)} \qquad \sum_{S^{(n_4)}}^{(n_3)} \qquad \sum_{\varphi^{(n_4)}}^{(n_3)} \qquad \sum_{\varphi^{(n_4)}}^{(n_3)} \qquad \sum_{\varphi^{(n_4)}}^{(n_3)} \qquad \sum_{\varphi^{(n_4)}}^{(n_4)} \qquad \sum_{\varphi^{(n_4)}}^{(n_4)}$$

$$\mathcal{M}\left[S^{(n_1)}S^{(n_2)} \to \varphi^{(n_3)}\varphi^{(n_4)}\right] = \frac{\kappa^2 s}{32} (1 - \cos 2\theta) \left\langle k^{(n)}k \right\rangle$$

t extra dimensions  $\varphi^{(n_4)}$ 



$$\mathcal{L}_m = \sqrt{G} \left( \frac{1}{2} G^{MN} \partial_M S \partial_N S - \frac{1}{2} M_S^2 S^2 \right)$$

 $k^{(n)}f^{(n)}_Sf^{(n)}_S
angle + \mathcal{O}(s^0)$  Agrees with Unitary gauge calculation





Expectation from Goldstone Equivalence Theorem

$$\mathcal{M}\left[h_L^{(n_1)}h_L^{(n_2)} \to h_L^{(n_3)}h_L^{(n_4)}\right] = \mathcal{M}\left[\varphi^{(n_1)}\varphi^{(n_2)} \to \varphi^{(n_2)}\right]$$

Agrees with Unitary gauge calculation

$$\mathcal{M}\left[\varphi^{(n_1)}\varphi^{(n_2)} \to \varphi^{(n_3)}\varphi^{(n_4)}\right] = \frac{\kappa^2 s}{64} \frac{(\cos 2\theta + 7)^2}{\sin^2 \theta}$$

 $\left|\varphi^{(n_3)}\varphi^{(n_4)}\right| + \mathcal{O}(s^0)$ 

 $\langle k^{(n)}k^{(n)}k^{(n)}k^{(n)}\rangle + \mathcal{O}(s^0)$ 

# A Short Summary of Double Copy: BCJ double copy

# A Short Summary of Double Copy: BCJ double copy

### Consider Gluon Scattering Amplitudes



$$n_{s} = -\frac{1}{2} \left\{ \left[ (\varepsilon_{1} \cdot \varepsilon_{2}) p_{1}^{\mu} + 2(\varepsilon_{1} \cdot p_{2}) \varepsilon_{2}^{\mu} - (1 \leftrightarrow 2) \right] \left[ (\varepsilon_{3} \cdot \varepsilon_{4}) p_{3\mu} + 2(\varepsilon_{3} + s \left[ (\varepsilon_{1} \cdot \varepsilon_{3}) (\varepsilon_{2} \cdot \varepsilon_{4}) - (\varepsilon_{1} \cdot \varepsilon_{4}) (\varepsilon_{2} \cdot \varepsilon_{3}) \right] \right\},$$

### The Contact Interaction is factored into the above definition by suitable kinematic reshuffling

# A Short Summary of Double Copy: BCJ double copy

### Consider Gluon Scattering Amplitudes



$$n_{s} = -\frac{1}{2} \left\{ \left[ (\varepsilon_{1} \cdot \varepsilon_{2}) p_{1}^{\mu} + 2(\varepsilon_{1} \cdot p_{2}) \varepsilon_{2}^{\mu} - (1 \leftrightarrow 2) \right] \left[ (\varepsilon_{3} \cdot \varepsilon_{4}) p_{3\mu} + 2(\varepsilon_{3} + s \left[ (\varepsilon_{1} \cdot \varepsilon_{3}) (\varepsilon_{2} \cdot \varepsilon_{4}) - (\varepsilon_{1} \cdot \varepsilon_{4}) (\varepsilon_{2} \cdot \varepsilon_{3}) \right] \right\},$$

## The Contact Interaction is factored into the above definition by suitable kinematic reshuffling

### Gauge Invariance demands that the amplitude must vanish under polarization to momentum replacement

$$n_{s}|_{\varepsilon_{4} \to p_{4}} = -\frac{s}{2} \Big[ (\varepsilon_{1} \cdot \varepsilon_{2}) \big( (\varepsilon_{3} \cdot p_{2}) - (\varepsilon_{3} \cdot p_{1}) \big) + \operatorname{cyclic}(1, 2, 3) \Big] \equiv s \, \alpha(\varepsilon, p) \qquad \left| \begin{array}{c} \frac{n_{s}c_{s}}{s} + \frac{n_{t}c_{t}}{t} + \frac{n_{u}c_{u}}{u} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \Big|_{\varepsilon_{4} \to p_{4}} \Big|_{\varepsilon_{4$$

$$n_{s}\big|_{\varepsilon_{4} \to p_{4}} = -\frac{s}{2}\Big[(\varepsilon_{1} \cdot \varepsilon_{2})\big((\varepsilon_{3} \cdot p_{2}) - (\varepsilon_{3} \cdot p_{1})\big) + \operatorname{cyclic}(1, 2, 3)\Big] \equiv s \,\alpha(\varepsilon, p) \qquad \left| \begin{array}{c} \frac{n_{s}c_{s}}{s} + \frac{n_{t}c_{t}}{t} + \frac{n_{u}c_{u}}{u} \Big|_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t} + c_{u}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t} + c_{u}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t}}{u} \right]_{\varepsilon_{4} \to p_{4}} = (c_{s} + c_{t}) + \frac{1}{2} \left[ \frac{c_{s} + c_{t}}$$





Color and Kinematic Factors are mutually Interchangable





Color and Kinematic Factors are mutually Interchangable

$$i\mathcal{A}_4^{\text{tree}} = g^2 \left( \frac{n_s c_s}{s} + \frac{n_t c_t}{t} + \frac{n_u c_u}{u} \right)$$

Massless Gravity amplitude





Color and Kinematic Factors are mutually Interchangable

$$\begin{split} i\mathcal{A}_{4}^{\text{tree}} &= g^{2} \bigg( \frac{n_{s}c_{s}}{s} + \frac{n_{t}c_{t}}{t} + \frac{n_{u}c_{u}}{u} \bigg) \\ \\ & \\ \text{Massless Gravity amplitude} \\ \hline i\mathcal{A}_{4}^{\text{tree}} \Big|_{\substack{c_{i} \to \tilde{n}_{i} \\ g \to \kappa/2}} &\equiv i\mathcal{M}_{4}^{\text{tree}} = \bigg( \frac{\kappa}{2} \bigg)^{2} \bigg( \frac{n_{s}^{2}}{s} + \frac{n_{t}^{2}}{t} + \frac{n_{u}^{2}}{u} \bigg) \\ \\ & \\ & \\ \text{Diffeomorphism Invariance} \\ \hline \frac{n_{s}^{2}}{s} + \frac{n_{t}^{2}}{t} + \frac{n_{u}^{2}}{u} \Big|_{\varepsilon_{4}^{\mu\nu} \to p_{4}^{\mu} \varepsilon_{4}^{\nu} + p_{4}^{\nu} \varepsilon_{4}^{\mu}} = 2(n_{s} + n_{t} + n_{u}) \,\alpha(\varepsilon, p) = 0 \,. \end{split}$$









Calculate the 4 D KK amplitudes of a Gauge theory

Calculate the 4 D KK amplitudes of a Gauge theory

 $\left| \hat{A}_0^a(\hat{p}_1) \hat{A}_0^b(\hat{p}_2) \to \hat{A}_0^c(\hat{p}_3) \hat{A}_0^d(\hat{p}_4) \right|$ 

 $\left| \hat{p}^M = (p^\mu, \pm p^5) \right|$ 

**On** -Shell Conditions

$$p_i^{\mu} p_{i\mu} = E^2 \cos^2 \omega \equiv M_n^2$$

## Calculate the 4 D KK amplitudes of a Gauge theory

 $\hat{A}_0^a(\hat{p}_1)\hat{A}_0^b(\hat{p}_2) \to \hat{A}_0^c(\hat{p}_3)\hat{A}_0^d(\hat{p}_4)$ 

$$\hat{p}^M = (p^\mu, \pm p^5)$$

$$\mathcal{M}_{5} \equiv \frac{1}{4} \sum_{\eta_{i}=\pm} \mathcal{M} \left[ \hat{A}_{0}^{a}(\hat{p}_{1,\eta_{1}}) \hat{A}_{0}^{b}(\hat{p}_{2,\eta_{2}}) \to \hat{A}_{0}^{c}(\hat{p}_{3,\eta_{3}}) \hat{A}_{0}^{d}(\hat{p}_{4,\eta_{4}}) \right] \delta_{\eta_{1}+\eta_{2}-\eta_{3}-\eta_{4},0} =$$

**On** -Shell Conditions

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$$\frac{1}{g_5^2}\mathcal{M}_5 - \frac{1}{g^2}\mathcal{M}_4 = -\frac{4c\left(\kappa^4 + \kappa^2 - 2\right)\left(C^{abe}C^{cde} + C^{ace}C^{dbe}\right)}{\kappa^2}$$

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 $^{e} + C^{ade}C^{bce}$ 



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$$\frac{1}{g_5^2}\mathcal{M}_5 - \frac{1}{g^2}\mathcal{M}_4 = -\frac{4c\left(\kappa^4 + \kappa^2 - 2\right)\left(C^{abe}C^{cde} + C^{ace}C^{dbe}\right)}{\kappa^2}$$

Combine two copies of different helicities and replace color factors by kinematic functions

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Combine two copies of different helicities and replace color factors by kinematic functions

$$\mathbf{M} = \left(\frac{\kappa}{2}\right)^2 \left(\frac{n_s n'_s}{s} + \frac{n_t n'_t}{t} + \frac{n_u n'_u}{u}\right)$$

### On - Shell Conditions

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lent to a 4D KK graviton amplitude on a flat torus





- Compactified theories of extra dimensions -> No low energy cut-off
- Cancellations due to different diagrams reduce  $O(s^5)$  growth to O(s).
- No low energy cut-off for consistent models of stabilization
- Uncovered sum rules enforcing this cancellation
- Can show -> Analysis extends to matter on brane or bulk
- Consistent with literature on massive gravity.
- Pheno papers : Doing an unitarity analysis for DM models, ultralight radion as a candidate ...
- Theory papers : Spinor Helicity/Goldstone Equivalence calculation.
- More connections with massive gravity community ...

• Possible to double-copy a compactified gauge theory to compactified gravity for flat toroidal compactification

